



NASA GSFC mascons: Advancements in signal covariance design and the estimation of noise uncertainties and leakage errors

B.D. Loomis¹, S.B. Luthcke¹, T. Sabaka¹, K. Rachlin^{1,2}

¹ NASA GSFC, Geodesy and Geophysics Laboratory, Greenbelt, MD

² SGT Inc. at NASA GSFC, Greenbelt, MD

GRACE Science Team Meeting

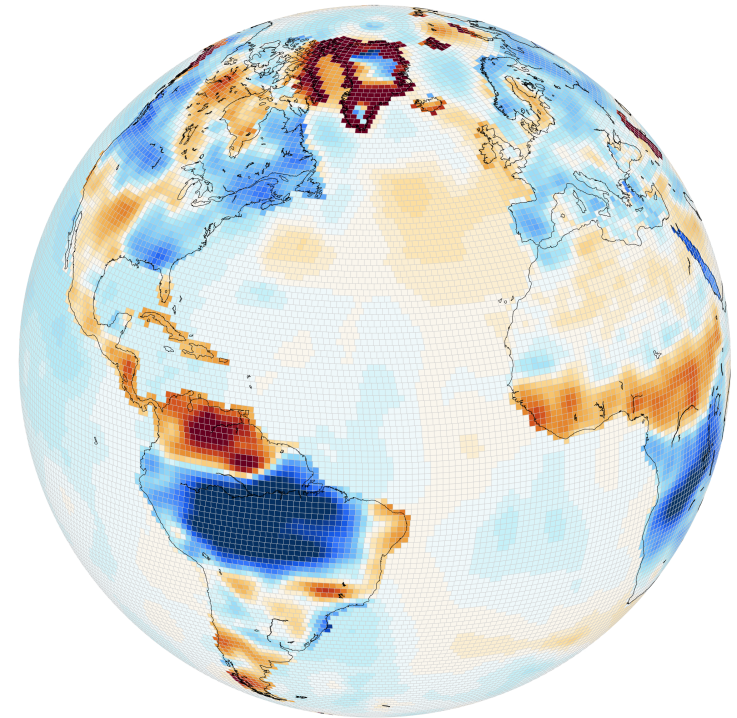
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Outline

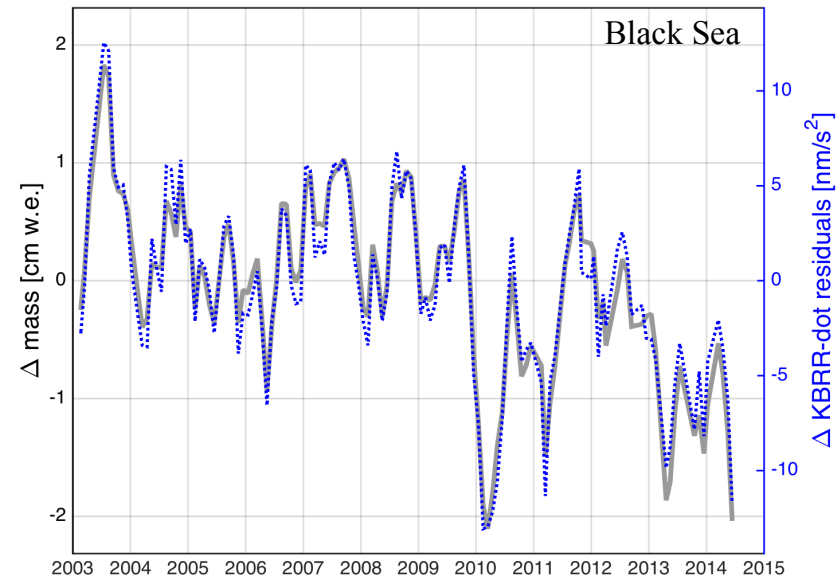
- GSFC mascon solution:
 - New constraint matrices
- GSFC mascon errors:
 - Noise uncertainties
 - Leakage errors
 - Error combination & validation



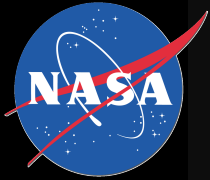


GSFC solution: New constraint matrices

- ❑ [Loomis and Luthcke, 2016] compared GSFC mascons to steric-corrected altimetry in large inland seas
- ❑ Demonstrated linear relationship between Δ local mass anomaly and Δ range-acceleration residuals, $\ddot{\rho}$ (once long-wavelength signals are recovered)



- ❑ New GSFC solution (v2.3b) applies local $\ddot{\rho}$ residuals to construct constraint matrices
- ❑ The same exponential taper constraints as [Luthcke et al., 2013] are still applied, except we now use mascon-dependent weighting



GSFC solution: New constraint matrices

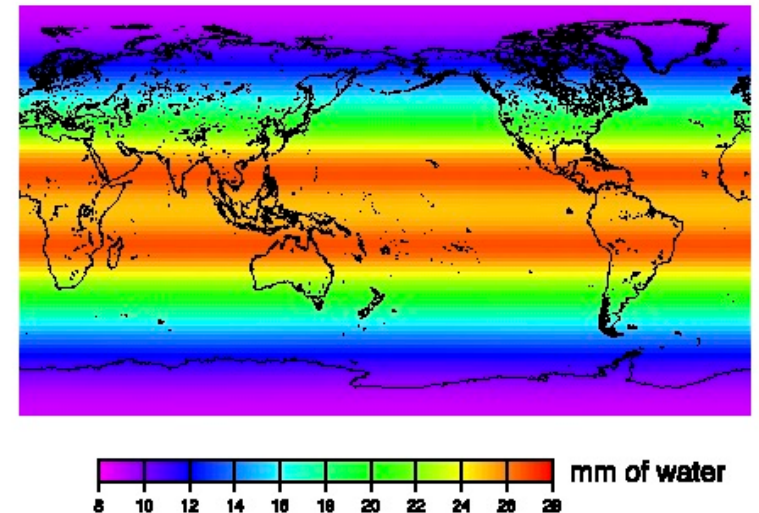
Iter.	TVG forward model	Constraints	Output
1	Trend & annual	Static; latitude-dependent	$\Delta\hat{x}_1$
2	Trend & annual + $\Delta\hat{x}_1$	Time-dependent; $\ddot{\rho}$ residuals	$\Delta\hat{x}_2$
3	Trend & annual + $\Delta\hat{x}_1 + \Delta\hat{x}_2$	Time-dependent; $\ddot{\rho}$ residuals	$\Delta\hat{x}_3$
4	Trend & annual + $\Delta\hat{x}_1 + \Delta\hat{x}_2 + \Delta\hat{x}_3$	N/A	Post-fits

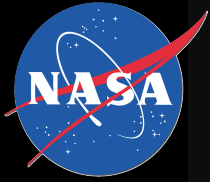


GSFC solution: New constraint matrices

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- *Static constraint matrix* with latitude-dependence, following error analysis of [Wahr et al., 2006]
- Relatively strong damping
- Stronger constraints in ocean due to low SNR
- By design the 1st mascon update, $\Delta\hat{x}_1$, only recovers large spatial scales

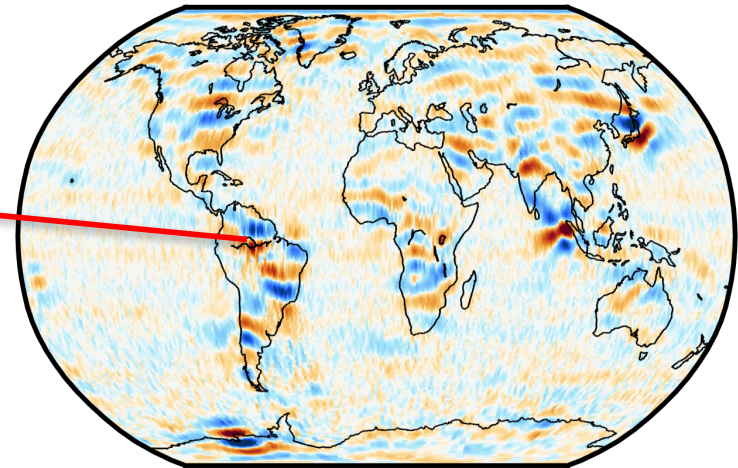
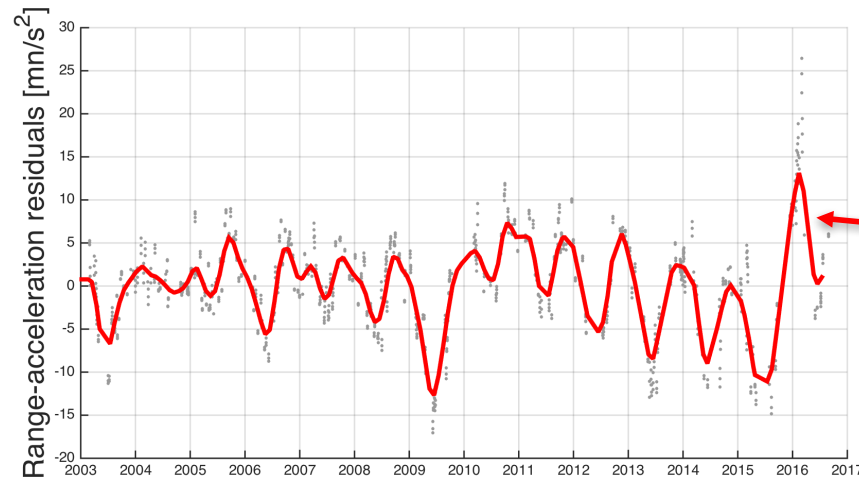




GSFC solution: New constraint matrices

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- *Time-dependent constraint matrices* from mascon-binned range-acceleration residuals
- 2nd and 3rd mascon updates, $\Delta\hat{x}_2$ & $\Delta\hat{x}_3$, recover TVG at smaller spatial scales

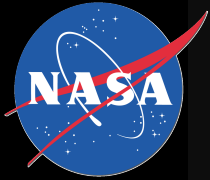




GSFC solution: New constraint matrices

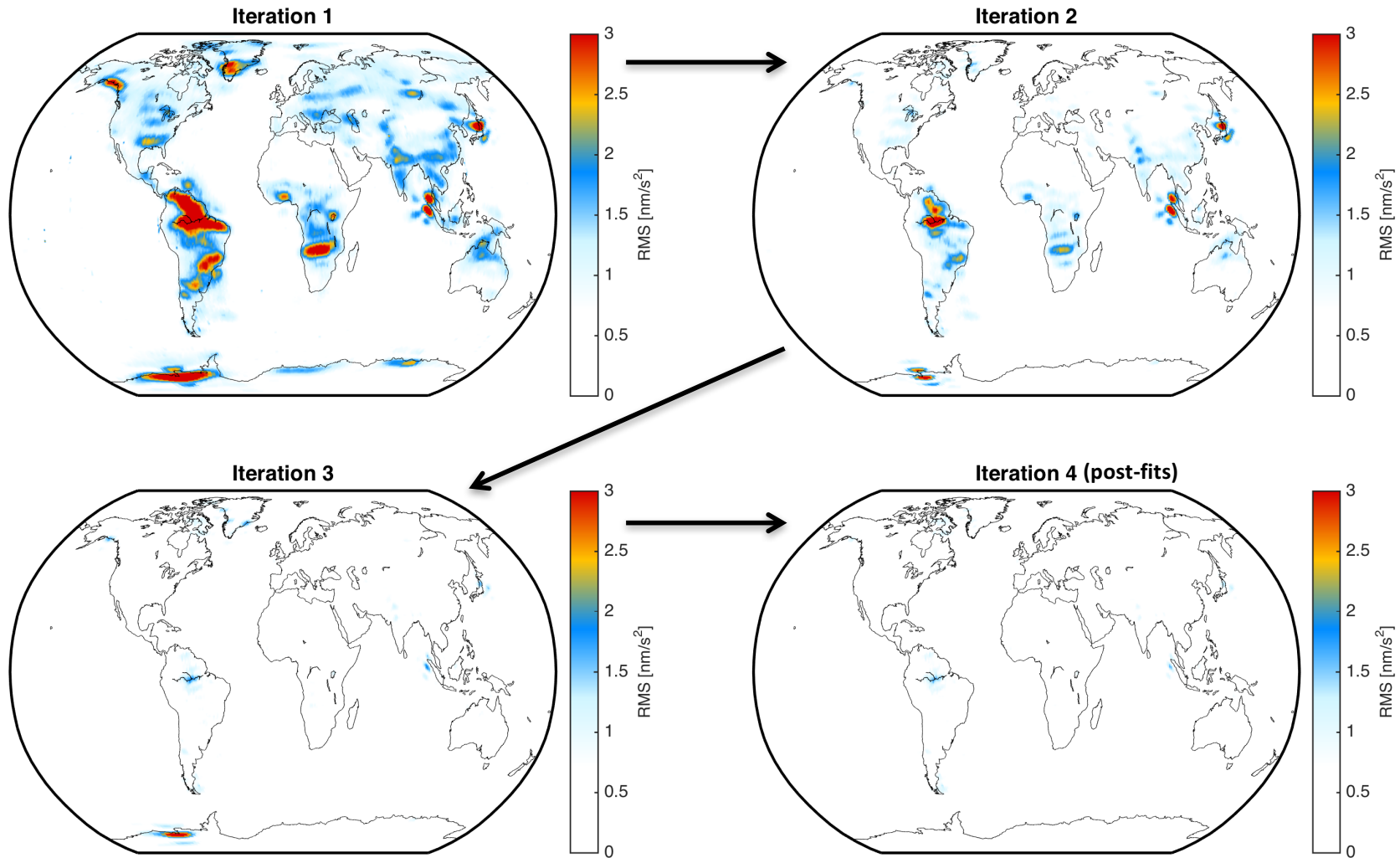
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- Final solution = Trend & annual + $\Delta\hat{x}_1 + \Delta\hat{x}_2 + \Delta\hat{x}_3$
- Post-fit KBRR residuals confirm the solution has converged and validate the estimated solution uncertainties



GSFC solution: New constraint matrices

- Reduction of range-acceleration residuals with each iteration





GSFC errors: Noise uncertainties

- ❑ One possible approach is to compute the actual monthly covariance matrices, but these require calibration, and are highly sensitive to the constraint matrix, P_m :

$$\text{Cov}(\hat{x}_i) = (A^T W A + \lambda P_m)_i^{-1}$$

- ❑ General definition of covariance matrix:

$$\text{Cov}(\hat{x}) \equiv \mathbb{E}([\hat{x} - \mathbb{E}(\hat{x})] [\hat{x} - \mathbb{E}(\hat{x})]^T)$$

- ❑ We define the expected value of the solution as the temporally filtered solution:

$$\mathbb{E}(\hat{x}) \cong \mathcal{F}(\hat{x}), \text{ where } \mathcal{F} \text{ is a tuned Savitzky-Golay filter}$$

- ❑ We now define the $M \times N$ noise matrix, \hat{n} (*no. mascons* \times *no. months*):

$$\hat{n} \equiv \hat{x} - \mathcal{F}(\hat{x})$$

- ❑ The spatial and temporal covariance matrices are numerically computed as:

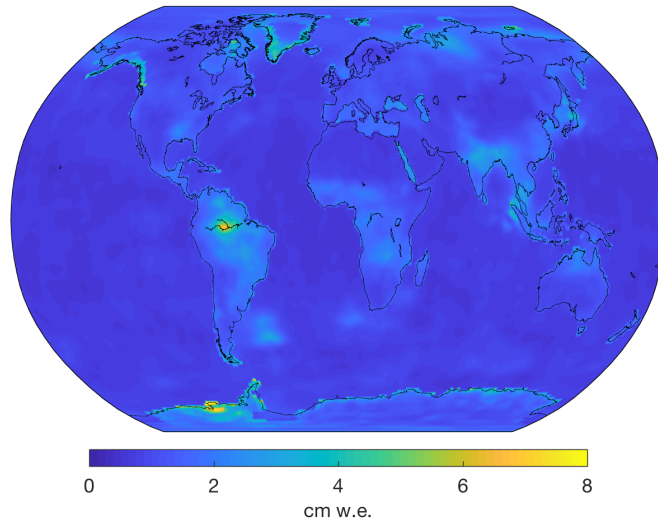
$$\text{Spatial: } \text{Cov}(\hat{x}) = \frac{1}{N-1} \sum_{i=1}^N \hat{n} \hat{n}^T$$

$$\text{Temporal: } \text{Cov}(\hat{x}) = \frac{1}{M-1} \sum_{i=1}^M \hat{n}^T \hat{n}$$

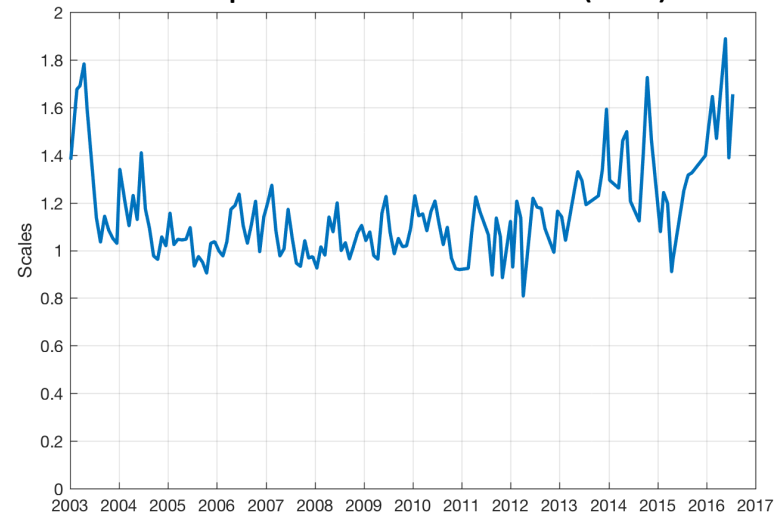


GSFC errors: Noise uncertainties

Spatial uncertainties ($1-\sigma$)



Temporal uncertainties ($1-\sigma$)



Mean of monthly $1-\sigma$ uncertainties [cm w.e.]

Greenland Ice Sheet	2.92
Antarctic Ice Sheet	2.25
Land	1.63
Ocean	0.88

- Uncertainties were successfully validated against two other methods:
 - MAD of the finest scale wavelet coefficients [Donoho and Johnstone, 1994]
 - KBRR post-fit residuals applied to semi-analytic error analysis [Kim, 2000]



GSFC errors: Leakage

- ❑ Leakage is computed with resolution operator, R (e.g. Luthcke et al., 2013):

$$\hat{x} = Rx$$

$$\ell = (R - I)x$$

where,

$$R = (A^T W A + \lambda P_m)^{-1} A^T W A$$

x = True mascon state

\hat{x} = Estimated mascon state (GSFC solution)

- ❑ Of course, the true mascon state is unknown, so we approximate the leakage error using our own solution, \hat{x} :

$$\hat{\ell} \approx (R - I)\hat{x}$$

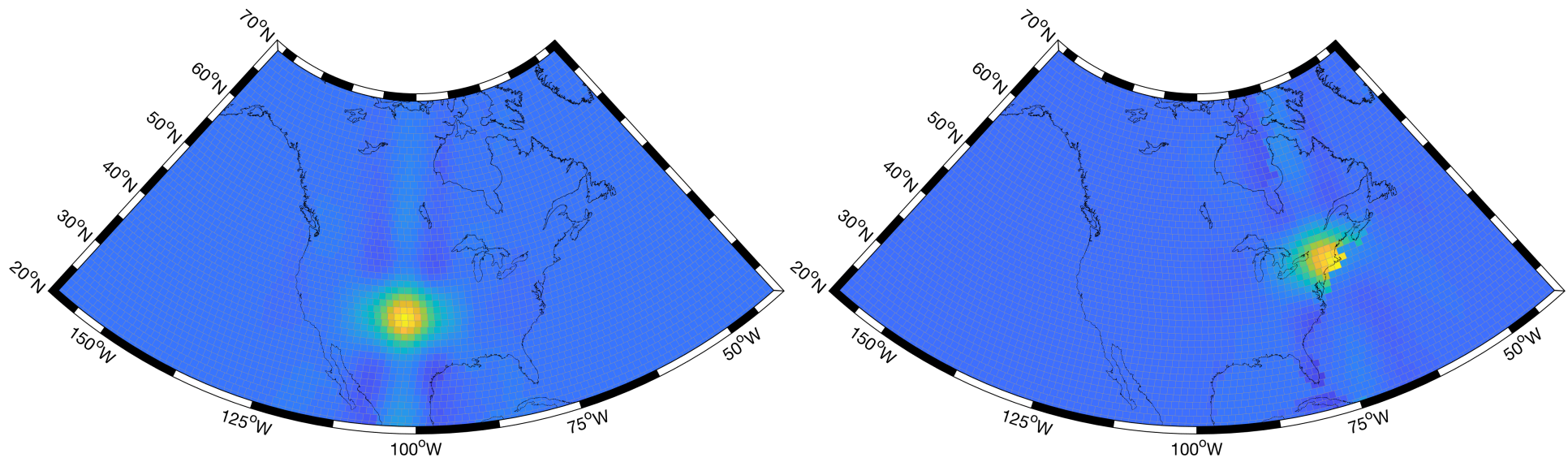
- ❑ Full $41,168 \times 41,168$ resolution operators are computed monthly, and applied to each monthly solution to produce monthly maps of leakage error



GSFC errors: Leakage

- Properties of resolution operator:
 - Rows of R : weighted averages; i.e. point-spread function
 - Columns of R : cross-talk between mascons
- Mascon resolution can be quantified with impulse response to R :

$\hat{x}_i = Rx_i$, where x_i has one non-zero mascon:





GSFC errors: Leakage

- Application of the monthly leakage maps
 - The leakage has clear deterministic (trend) and stochastic components:

$$\ell_{total} = \ell_{trend} + \ell_{random}$$

$$\ell_{\sigma} \equiv STDEV(\ell_{random})$$

$$\ell_{95\%} = \ell_{trend} + 2\ell_{\sigma}$$

- The leakage maps are not a correction to the solution, but rather an estimate of the leakage error



GSFC errors: Combination & validation

- Combining errors:

$$MSE = bias^2 + \sigma^2$$

- Combining errors for single mascon:

$$95\% \text{ uncertainty} = \ell_{trend} + 2\ell_{\sigma} + 2\sigma_{noise}$$

- Combining errors for region/basin:

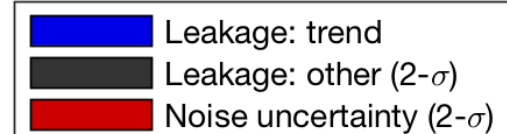
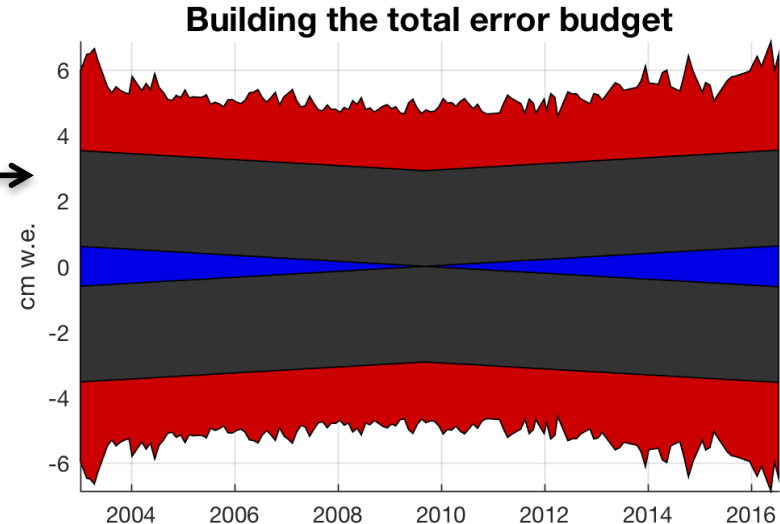
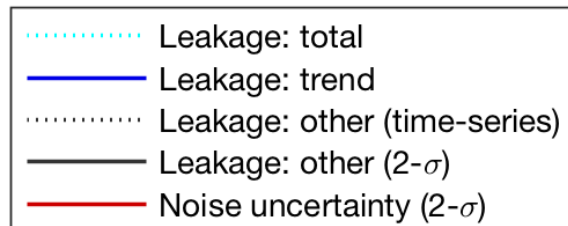
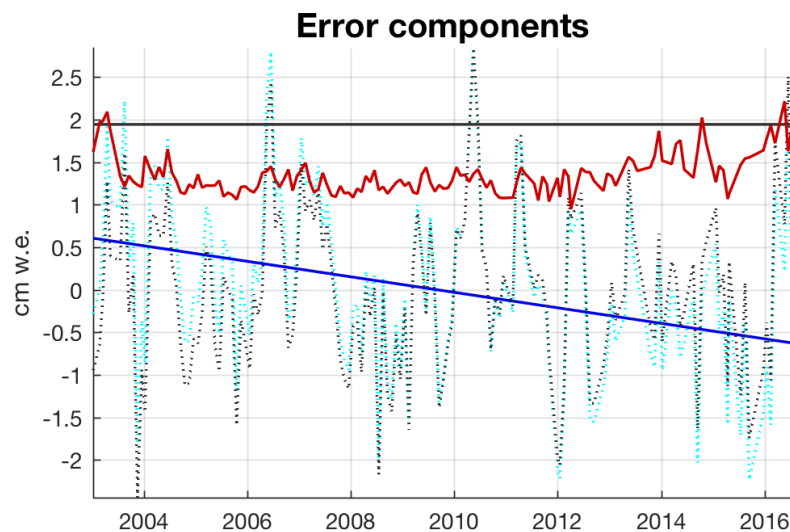
$$95\% \text{ uncertainty} = \overline{\ell_{trend}} + 2\overline{\ell_{\sigma}} + \frac{2\overline{\sigma_{noise}}}{\sqrt{\frac{N}{22}}}$$

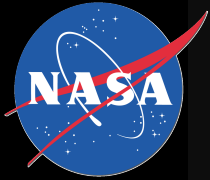
- $\overline{\ell_{trend}}$ and $\overline{\ell_{\sigma}}$ are computed from the regional leakage time series
- $\overline{\sigma_{noise}}$ is the average uncertainty over the region
- N is the number of mascons in the region
- $\sqrt{N/22}$ accounts for uncorrelated uncertainties
- 22 mascons is approximate spatial resolution of ~300 km
- if $N < 22$, just use $2\overline{\sigma_{noise}}$ as the uncertainties are correlated



GSFC errors: Combination & validation

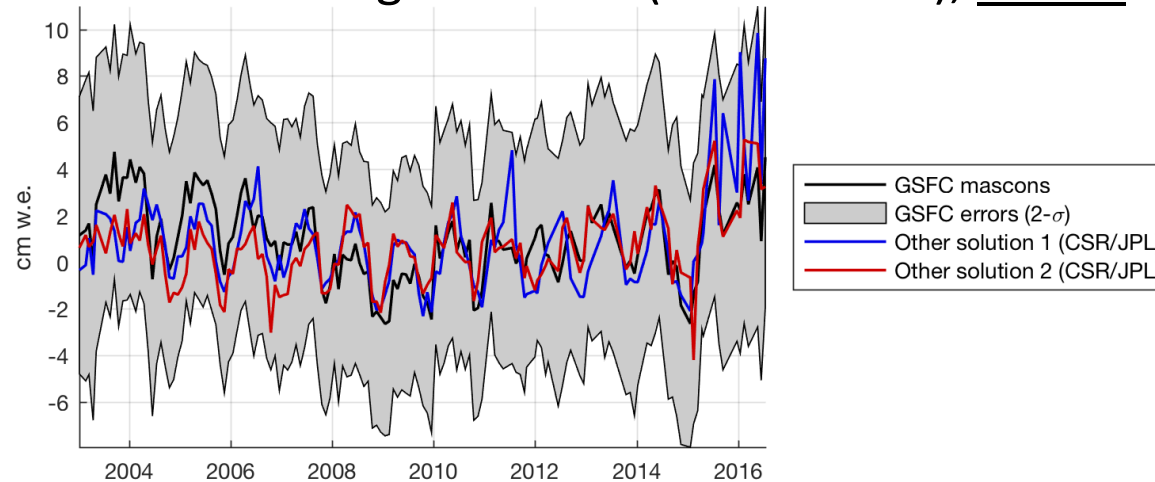
$$95\% \text{ uncertainty} = \ell_{trend} + 2\ell_{\sigma} + 2\sigma_{noise}$$



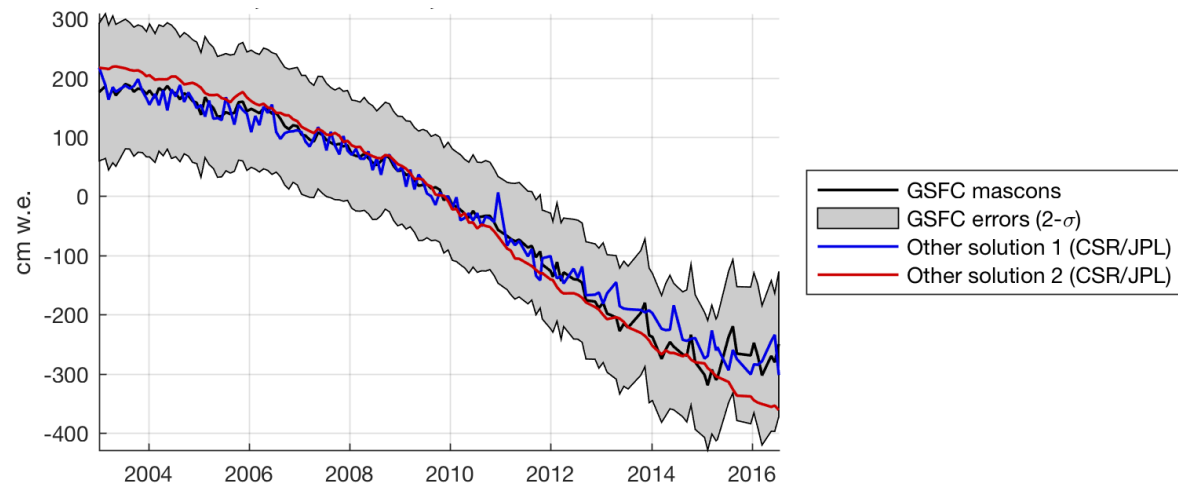


GSFC errors: Combination & validation

- Sample application for a single mascon (Central Asia), SNR<1



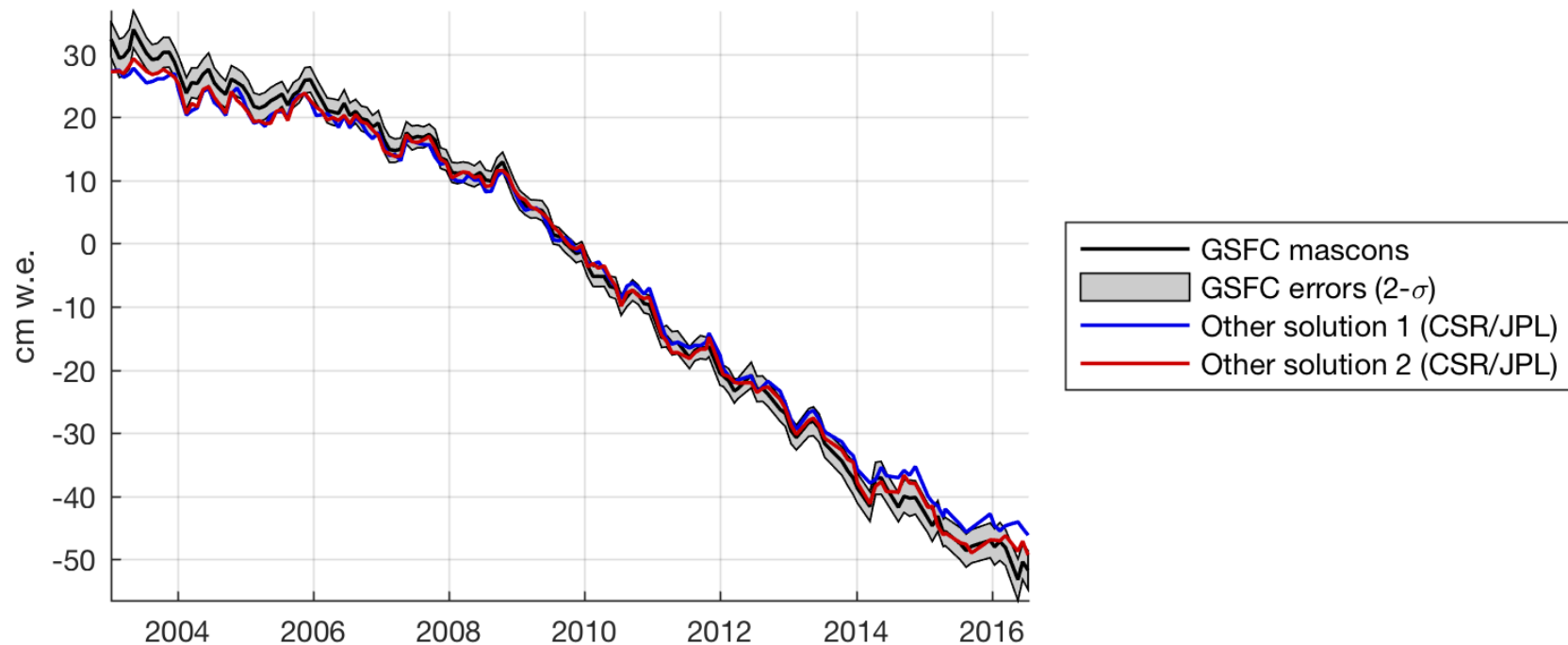
- Sample application for a single mascon (West Antarctic Ice Sheet), SNR>1





GSFC errors: Combination & validation

- ❑ Sample application for mascon region (West Antarctic Ice Sheet)



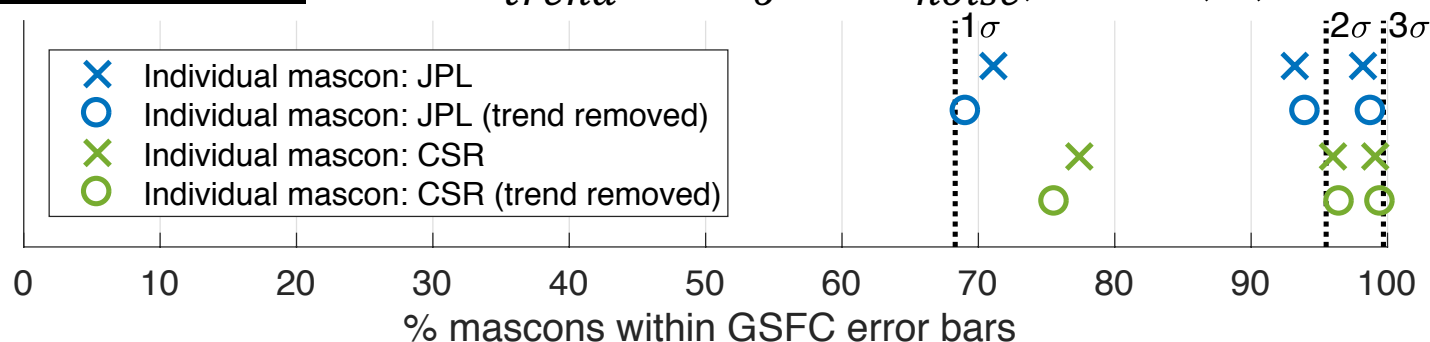


GSFC errors: Combination & validation

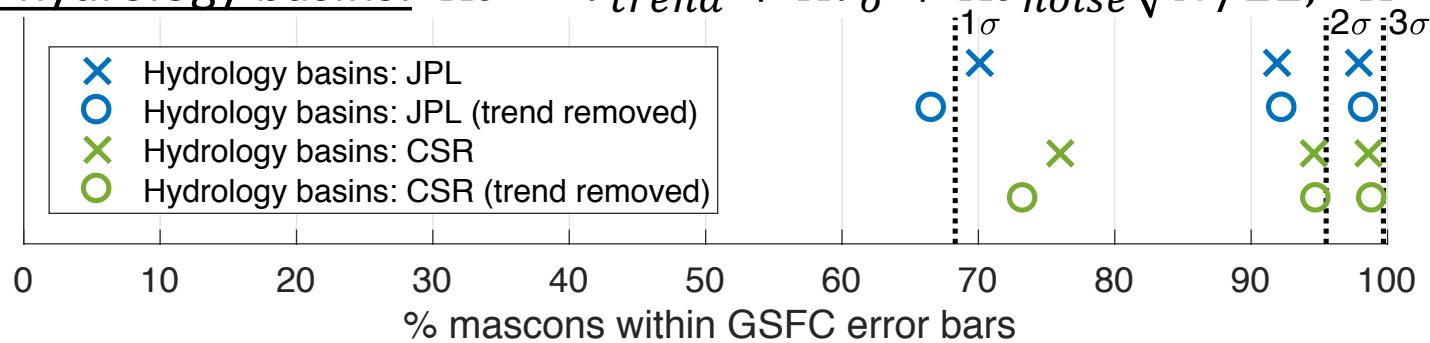
- How well do the GSFC error bars contain the JPL & CSR mascons?

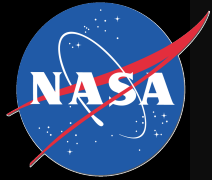
If the error bars are well-constructed, we would expect the 1- σ , 2- σ , 3- σ errors bars to contain ~68%, 95%, 99% of independent solution data points

- Individual mascons: $K\sigma = \ell_{trend} + K\ell_{\sigma} + K\sigma_{noise}$, $K = 1, 2, 3$



- Global hydrology basins: $K\sigma = \overline{\ell_{trend}} + K\overline{\ell_{\sigma}} + K\overline{\sigma_{noise}}\sqrt{N/22}$, $K = 1, 2, 3$





Summary

- ❑ Latest GSFC mascon solution (v2.3b) applies range-acceleration residuals in design of iterative mascon constraints
- ❑ New mascon noise uncertainties validated against KBRR residuals
- ❑ Leakage errors determined with full monthly resolution operators
- ❑ Resolution operator provides clear way to define spatial resolution of solution, however, instead of just considering mascons in terms of their resolution (e.g. 150,000 km²), it is better to look at signal-to-noise ratio
 - There are individual 1-arc-degree mascons with $\text{SNR} > 1$
 - There are basins larger than GRACE “resolution” with $\text{SNR} < 1$
- ❑ Upcoming solution release will have updated noise uncertainties, leakage error, and recipe to properly combine them
- ❑ NASA GSFC mascon website: <https://neptune.gsfc.nasa.gov/grace>